New States of Quantum Matter

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Recent Progress in Many-Body Theories 14 Barcelona



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New states of quantum matter created in the past decade

From:

Trapped cold atomic systems: Bose-condensed and BCS fermion superfluid states T ~ nanokelvin (traps are the coldest places in the universe!)

To:

Deconfined quark-gluon plasmas made in ultrarelativistic heavy ion collisions $T \sim 10^2 \text{ MeV} \sim 10^{12} \text{ K}$ (temperature of early universe at $\sim 1 \mu$ sec)

Separated by ~21 decades in characteristic energy scales, yet have intriguing overlaps.

Cold atoms: trapped bosons and fermions



Box



Potential well (trap)

Statistics:

Bose condensate: macroscopic occupation of single mode (generally lowest)





Degenerate Fermi gas



=> BCS pairing

Trapped atomic experiments in a nutshell

Warm atomic vapor



T=300K, $n \sim 3 \times 10^6$ /cm³

Magneto-optical trap



 $\begin{array}{l} \text{Laser cool to } T\sim 50 \mu \text{K} \\ n\sim 10^{11} \text{/cm}^3 \end{array}$

Evaporatively cool in magnetic (or optical) trap



Bosons condense, Fermions BCS-pair $T \sim 1-10^3$ nK $n \sim 10^{14-15}$ /cm³ $N \sim 10^5$ -10⁸

Experiment, and then measure :

To probe system, release from trap, let expand and then image with laser:



Long-Lived Alkali Atoms			
BOSONS (Spin, lifetime) (Z-N=odd-even nuclei)		FERMIONS (Z-N=odd-odd nuclei)	
⁷ Li 3/2-		⁶ Li 1+	
²³ Na 3/2-		²² Na 3+	2.6y
³⁹ K 3/2+		⁴⁰ K 4-	1.3x10 ⁹ y
⁴¹ K 3/2+			
⁸⁵ Rb 5/2-		⁸⁶ Rb 2-	18.6d
⁸⁷ Rb 3/2-	4.75x10 ¹⁰ y		
¹³¹ Cs 5/2+	9.7d	¹³² Cs 2+	6.5d
¹³³ Cs 7/2+			
¹³⁵ Cs 7/2+	2.3x10 ⁶ y	¹³⁴ Cs 4+	2.06y
²⁰⁹ Fr 9/2-	50.0s	²⁰⁸ Fr 7+	59.1s

Early days of ultracold trapped atomic gases \geq 1995 = first Bose condensation of ⁸⁷Rb, ²³Na and ⁷Li

*Structure of condensate.



*Elementary modes: breathing, quadrupole, short wave sound, ...



*1, 2 and 3 body correlations => evidence for BEC rather than simply condensation in space.

*Interference of condensates.



Primarily described in terms of mean field theory – Gross-Pitaevskii eq. $i\hbar\partial \psi(\mathbf{r},t) /\partial t = [-\hbar^2 \nabla^2 / 2m + V(\mathbf{r}) + g|\psi(\mathbf{r},t)|^2]\psi(\mathbf{r},t)$

Recent directions in ultracold atomic systems, I Strongly correlated systems

* Rapidly rotating bosons: how do many-particle Bose systems carry extreme amounts of angular momentum?

•Trapping and cooling clouds of fermionic atoms Degenerate Fermi gases and molecular states BCS pairing => new superfluid Crossover from BEC of molecules to BCS paired state

* Physics in the strong interaction limit: scale-free regime where $r_0 \ll n^{-1/3} \lesssim a$ $r_0 =$ range of interatomic potential \sim few Å n = particle density a = s-wave scattering length Realize through atomic Feshbach resonances

Recent directions in ultracold atomic systems, II Novel systems

*Physics in optical lattices: Mott transition from superfluid to insulating states; low dimensional systems; 2D superfluids

* Spinor gases: trapped by laser fields. Physics of spin degrees of freedom Fragmented condensates

* Mixtures of bosons and fermions

 * Ultracold molecules: coherent mixtures of atoms and molecules, e.g., ⁸⁷Rb atoms and ⁸⁷Rb₂ molecules; heteronuclear molecules: ⁶Li+²³Na, ⁴⁰K+⁸⁷Rb

Future applications:

Trapped ions for quantum computing Slow light

Slow light

Atom lithography

Matter lasers

Laboratories for ultracold physics appropriately situated

JILA (Boulder, CO)

Vortices in trapped atomic clouds

Illinois, every spring

Vortices in superfluids

Spin container of superfluid (e.g., helium) slowly. Liquid remains at rest

Spin fast enough. Form vortex in center of liquid!

Superfluid ⁴He viewed along rotation k axis. Imaged by trapping electrons in cores *Yarmchuk, Gordon, & Packard, PRL43, 214 (1979)*

Making vortices in Bose-Einstein condensates

Bose condensed ⁸⁷Rb (ENS)

K. W. Madison, F. Chevy, W. Wohlleben, J. Dalibard 1999

Rapidly rotating superfluid contains triangular lattice of vortices

Abo-Shaeer et al. (MIT) 2001

Engels et al. (JILA) 2002

As Ω grows in harmonic trap, vortex lattice melts, and go through a sequence of new highly correlated states with large angular momentum, L/N ~ (10² - N) \hbar not yet reached experimentally.

Compress matter to form new states

Atoms

Plasma

Nuclear matter

$\rho \sim 2.5 \times 10^{14} \text{gm/cm}^3 = \rho_{\text{nm}} = 0.17 \text{ baryons/fm}^3$

 $fm = 10^{-13} cm$

Nucleons

Quark matter

Quark degrees of freedom

Quarks = fractionally charged spin-1/2 fermions, baryon no. = 1/3, with internal SU(3) color degree of freedom. {3 repr. of SU(3)}

Flavor	Charge/ e	Mass(MeV)
u	2/3	5 (2. <i>1-3.5</i>)*
d	-1/3	10 (2.1-3.5)*
S	-1/3	150 (54-92)*
С	2/3	1300
b	-1/3	4200
t	2/3	175000

Hadrons are composed of quarks: proton = u + u + dneutron = u + d + d $\pi^+ = u + d$, etc.

*Lattice gauge theory calculations, Gough et al., PRL 79, 1622 (1997)

Form of baryons in the early universe at t < 1μ sec (T > 100 MeV). Possible basic degrees of freedom in deep interiors of neutron stars. PHASE DIAGRAM OF NUCLEAR MATTER

Quark-gluon plasma state

Degrees of freedom are deconfined quarks and gluons

Many more degrees of freedom than hadronic matter (color, spin, particle-antiparticle, & flavor); much larger entropy at given temperature.

<= Large latent heat (or sharp rise at least)

At low temperatures form Fermi seas of degenerate u,d, and s quarks: (e.g., in neutron stars?)

T.D. Lee

Creating high energy density matter in the lab

Relativistic Heavy Ion Collider (Brookhaven) since 2000. Colliding beams 100 GeV/ALarge Hadron Collider (CERN) in 2008.2700 GeV/A

<u>100 GeV per nucleon</u> Au(197 × 100) + Au(197 × 100)

What collisions actually look like in the lab. STAR detector

Schematic collision: Two Lorentz contracted nuclei collide, pass through each other, leaving highly excited state of vacuum in between.

ALICE detector at LHC

Two major detectors at RHIC PHENIX STAR Two smaller detectors BRAHMS PHOBOS

A few crucial observations at RHIC:

Produce matter with energy densities $\sim 5 \text{ GeV/fm}^3$ $\sim 10-30 \times \text{energy}$ density of ordinary nuclei $\sim 0.15 \text{ GeV/fm}^3$

Certainly produce quark-gluon plasma.

Fast quarks traversing medium lose energy rapidly. "Opaque" medium

Very rapid build-up of pressure in collisions: Large collective flow, fast thermalization, large interaction cross sections.

Hydrodynamics => small viscosity

Common problems of cold atom physics and RHIC physics: Small clouds with many degrees of freedom $\sim 10^4 - 10^7$ Strongly interacting systems

Infrared (long wavelength) problems in qcd and condensed bosons.

Recent connections:

Crossover: BEC ⇔ BCS and hadron ⇔ quark-gluon plasma

Viscosity: heavy-ion elliptic flow \Leftrightarrow Fermi gases near unitarity

Superfluidity and pairing in unbalanced systems: trapped fermions ⇔ color superconductivity

Ultracold ionized atomic plasma physics

Strong interactions

In quark-gluon plasma,

$$\alpha_s(p) = \frac{g_s^2}{4\pi} = \frac{6\pi}{(33 - 2N_f)\ln(p/\Lambda)}$$

Even at GUT scale, 10^{15} GeV, $g_s \sim 1/2$ (cf. electrodynamics: $e^2/4\pi = 1/137 \Longrightarrow e^{-1/3}$)

QGP is always strongly interacting

In cold atoms, effective atom-atom interaction is short range and s-wave:

 $V(r_1-r_2) = (4\pi\hbar a/m) \delta (r_1-r_2)$

a = s-wave atom-atom scattering length. Cross section: $\sigma = 8\pi a^2$ Go from weakly repulsive to strongly repulsive to strongly attractive to weakly attractive by dialing external magnetic field through Feshbach resonance.

 $\Lambda \sim 150 \text{ MeV}$

Feshbach resonance in atom-atom scattering

Low energy scattering dominated by bound state closest to threshold

Adjusting magnetic field, B, causes level crossing and resonance, seen as divergence of s-wave scattering length, a:

$$a(B) = a_{bg} \left(1 - \frac{\Delta}{B - B_{Feshbach}} \right)$$

BCS paired fermions: a new superfluid

Produce trapped degenerate Fermi gases: ⁶Li, ⁴⁰K

Increase attractive interaction with Feshbach resonance

At resonance have "unitary regime": no length scale

Experiments: JILA, MIT, Duke, Innsbruck, ...

Both systems scale-free in strongly coupled regime

 $F_{qgp} \sim const n_{exc}^{4/3}$ $E_{cold atoms} \sim const n^{2/3}/m$

Only length-scale for cold atoms near resonance is density. No microscopic parameters enter equation of state

$$\frac{E}{N} = \frac{3}{5}E_F(1+\beta)$$

 β is universal parameter. No systematic expansion

Fixed Node Green's Function Monte Carlo, Carlson et al. (2003): $\beta = -0.56 \text{ to} - 0.58$ Diagrammatic. Perali, Pieri & Strinati (2004) $\beta = -0.545$

Experiment: Rice: -0.54(5), Duke: -0.26(7), ENS: -0.3, JILA: -0.4, Innsbruck: 0.68(1)

BEC-BCS crossover in Fermi systems

Continuously transform from molecules to Cooper pairs: D.M. Eagles (1969) A.J. Leggett, J. Phys. (Paris) C7, 19 (1980) P. Nozières and S. Schmitt-Rink, J. Low Temp Phys. 59, 195 (1985)

⁶Li

Phase diagram of cold fermions vs. interaction strength

Phase diagram of quark gluon plasma

Karsch & Laermann, hep-lat/0305025

New critical point in phase diagram: induced by chiral condensate – diquark pairing coupling via axial anomaly

Hatsuda, Tachibana, Yamamoto & GB, PRL 97, 122001 (2006)

Abuki, Itakura & Hatsuda, PRD65, 2002

In SU(2)_C : hadrons <=> 2 fermion molecules, paired deconfined phase <=> BCS paired fermions

Possible structure of crossover *Fukushima, hep-ph/0403091*

Viscosity in elliptic flow in heavy ion collisions and in Fermi gases near unitarity

Strong coupling leads to low first viscosity η , seen in expansion in both systems

Shear viscosity η :

$$F = \eta A v / d$$

Stress tensor

Stress tensor
$$T_{diss}^{ij} = \eta \left(\frac{\partial v_i}{\partial x^j} + \frac{\partial v_j}{\partial x^i} - \frac{2}{3} \delta_{ij} \nabla \cdot v \right) + \zeta \delta_{ij} \nabla \cdot v$$
First viscosity $\eta \sim \rho \bar{v}^2 \tau \sim \frac{1}{|M|^2}$ $\tau = \text{scattering time}$

Strong interactions => small η

Strongly coupled ⁶Li expansion

Free Expansion:

K. M. O'Hara, S. L. Hemmer, M. E. Gehm, S. R. Granade, and J. E. Thomas, Science Dec 13 2002: 2179

Turn off trap: cloud expands

Compare with expansion of weakly coupled system →

Strongly coupled ⁶Li expansion

Free Expansion:

100 µs

1000 µs

1500 µs

K. M. O'Hara, S. L. Hemmer, M. E. Gehm, S. R. Granade, and J. E. *Thomas. Science Dec 13 2002: 2179*

Turn off trap: cloud expands

Pressure gradient largest in narrow direction

Expands asymmetrically

Similar to elliptic flow in heavy ion collisions

Find equation of state by fitting expansion with ideal (zero viscosity) hydrodynamics

Viscosity extracted from radial breathing mode

Expt: J. Kinast, A. Turlapov, J.E. Thomas, PRL 94, 170404 (2005)

Theory: T. Schaefer, cond-mat/0701251.

Ratio of shear viscosity to entropy density (\hbar =1)

Temperature/ Fermi temperature

Collectivity: Elliptic flow in non-central collisions:

anisotropic in ϕ (= azimuthal angle in x,z plane)

Crab Nebula Hubble Space Telescope • Wide Field Planetary Camera 2

PRC96-22a · ST Scl OPO · May 30, 1996 · J. Hester and P. Scowen (AZ State Univ.) and NASA

Hydrodynamic predictions of $v_2(p_T)$

Elliptic flow => almost vanishing viscosity in quark-gluon plasma

From T. Hirano

Conjectured lower bound on ratio of first viscosity to entropy density, s:

 $\eta > \hbar s/4\pi$

Kovtun, Son, & Starinets, PRL 94, 111601 (2005)

Figure 1: The viscosity-entropy ratio for some common substances.

(Exact result in \mathcal{N} =4 supersymmetric Yang-Mills theory in large N_c)

$$\eta \sim n_t m v^2 \tau = n p \lambda, \quad s \sim n_t$$

 $n_t = no.$ of degrees of freedom producing viscosity $p = mv = mean \text{ particle momentum} > \hbar / (interparticle spacing)$ $\lambda = mean \text{ free path}$

Bound \Leftrightarrow mean free path > interparticle spacing

Strongly coupled systems approach viscosity lower bound

Cold fermions in normal state at unitarity: $\eta \sim n\hbar T/T_{f}, s \sim n T/T_{f} \Rightarrow \eta/s \sim \hbar$ *G. Bruun and <u>H. Smith, cond-mat/06012460</u>*

Lattice calculations of first viscosity in qcd:

Nakamura & Sakai, hep-lat/0510039

 $\begin{array}{l} \mbox{Perturbative qcd limit:} \\ \eta \sim T^3/(\alpha_s{}^2 \ln \alpha_s) \\ \eta/S \sim 1/\alpha_s{}^2 \ln \alpha_s \end{array}$

GB, Monien, Pethick & Ravenhall, PRL 64(1990)

Shear viscosity of Fermi gas at unitarity

G. Rupak & T. Schaefer, arXiv:0707.1520

G. M. Bruun & H. Smith, PRA 75, 043612 (2007).

Shear viscosity/ entropy density ratio vs. T/T_F

Color pairing in quark matter

Review: Rajagopal & Wilczek, hep-ph/0011333

Superfluidity

condensate of paired quarks => superfluid baryon density (n_s)

Color Meissner effects

transverse color fields screened on spatial scale ~ London penetration depth ~ $(\mu/g^2n_s)^{1/2}$

Two interesting phases:

Color-flavor locked (CFL) $(m_u = m_d = m_s)$ $\langle u \rangle = \langle l \rangle = \langle s \rangle = \langle s \rangle$

Superfluidity and pairing for unbalanced systems

Trapped atoms: change relative populations of two states by hand

QGP: balance of strange (s) quarks to light (u,d) depends on ratio of strange quark mass m_s to chemical potential μ (>0)

Vortices as marker of superfluidity (MIT)

No. of vortices vs. population imbalance

Color superconductor with $m_{strange} \neq m_{light}$

Decreasing pairing of strange quarks with increasing m_s *Alford, Kovaris & Rajagopal, hep-ph/0311286*

Phase diagram in Δ_{CFL} , m_s² plane Abuki, Kitazawa, & Kunihiro, PLB 615, 102 (2005)

In gapless phase for unbalanced color superconductors, Meissner screening length can be imaginary (superfluid mass density < 0) *M. Huang; M. Alford; and collaborators*

Proposed resolutions

*Phase separation. (Cf. neutron-rich nuclei with a neutron skin.)

*FFLO state with crystalline ordering.

*Gluon condensate

. . .

*Current carrying states with non-zero spatially dependent order parameter, $\sim e^{i k \cdot r}$ (*T. Schäfer, nucl-th/0602067*)

Experiments on ⁶Li with imbalanced populations of two hyperfine states, |1⟩ and |2⟩

MIT: Zwierlein et al., Science 311, 492 (2006); Nature 442, 54 (2006).

Rice: Partridge et al., Science 311, 503 (2006) cond-mat/0605581

Fill trap with $n_1 |1\rangle$ atoms, and $n_2 |2\rangle$ atoms, with $n_1 > n_2$.

Study spatial distribution, and existence of superfluidity for varying $n_1:n_2$.

Phase diagram of trapped imbalanced Fermi gases

K. B. Gubbels, M. W. J. Romans, and H. T. C. Stoof, cond-mat/0606330

Sarma: second order transition to normal phase with increasing radius with gapless superfluid near boundary

Phase separation: first order transition

Spatial separation of condensate and unpaired atoms

majority state minority state unpaired atoms

Rice

Phase separation:

BEC side: repulsion between atoms and molecules.

BCS side: quasiparticle energy gap expels unpaired atoms from condensate.

Axial radius of cloud vs. polarization

Spatial separation vs. polarization

Partridge, Li, Liao, Hulet, Haque & Stoof, cond/mat 0608455

 $\mathbf{P} = (\mathbf{N}_1 - \mathbf{N}_2) / (\mathbf{N}_1 + \mathbf{N}_2)$

Spatial distribution in trap

No evidence of spatial modulation expected in FFLO state

Critical imbalance vs. coupling strength

New quantum phase transition: limit of superfluidity, at $\delta\mu \simeq \Delta$

At unitarity $P_c = (N_1 - N_2)/(N_1 + N_2) = 70(5)\%$ (Zwielein et al, cond-mat/0605258)

Ultracold neutral atomic plasmas

Killian, Kulin, Bergeson, Orozco, Orzel, & Rolston, PRL 83, 4776 (1999),
Kulin, Killian, Bergeson, & Rolston, PRL85, 318 (2000),
Killian, Chen, Gupta, Laha, Martinez, Mickelson, Nagel, Saenz, & Simien,
Proc. 12th Int. Cong. on Plasma Phys., 2004, physics/0410019,
Roberts, Fertig, Lim, & Rolston, physics/0402041.

Produce by photoionizing trapped cold atomic gas., e.g., Xe, Sr. In Xe, reach $T_e = 0.1 - 10^3$ K, $T_{ion} = 10\mu$ K - 4mK, $n = 2 \times 10^9$ /cm³, $N \sim 2 \times 10^5$ Expand plasma to measure

Optical depth of an Sr plasma $N = 7 \times 10^7$, $n \sim 2 \times 10^{10}$ /cm⁻³

Strongly coupled plasmas: $\Gamma = E_{interaction} / E_{kinetic} >> 1$ Electrons in a metal

$$\begin{split} E_{int} &\sim e^2/r_0 \quad r_0 = interparticle \ spacing \ \sim \hbar \ /k_f \\ E_{ke} &\sim k_f^2/m \Longrightarrow \ \Gamma \sim e^2/\hbar \ v_f = \alpha_{eff} \end{split}$$

 $v_{f} \sim 10^{-2} - 10^{-3} c \implies \alpha_{eff} \sim 1 - 5$

Dusty interstellar plasmas

Laser-induced plasmas (NIF, GSI)

Quark-gluon plasmas

 $E_{int} \sim g^2 / r_0$, $r_0 \sim 1/T$, $E_{ke} \sim T \implies \Gamma \sim g^2 \gg 1$

Ultracold trapped atomic plasmas

 $\Gamma \sim n_9^{1/3}/T_K$ [where $n_9 = n/10^9$ /cm³ and $T_K = (T/1K)$] Non-degenerate plasma, $E_{ke} \sim T => \Gamma = E_{int}/E_{ke} \sim e^2/r_0T$ Ultracold plasmas analog systems for gaining understanding of plasma properties relevant to heavy-ion collisions:

- -kinetic energy distributions of electrons and ions -modes of plasmas: plasma oscillations -screening in plasmas -nature of expansion – flow, hydrodynamical (?) -thermalization times -correlations -interaction with fast particles
- -viscosity

Evolution of plasma temperatures

Ion temperature vs. time T. C. Killian et al., physics/0410019

Thermal equilibration on times $<< 1/\omega_{plasma}$ At short times, release of correlation energy heats the ions.

Electron temp. vs. time (Xe) J.L. Roberts et al., physics/0402041